## Basic concepts of modern standard cosmology: The three cosmic scalars!

- The universe is a spacelike homgeneity
- Cosmic matter density varies inversely proportional to the spacelike volume!
- The universe is characterized by a globally isotropic curvature; i.e. K=0; or =(+/-)1!
- The cosmic vacuum energy density is constant!

Then ---->>>

### Present-day standard cosmology with K=Lambda=const!

#### Kosmologische Konstante

Die Einstein-Gleichung kann formal einen weiteren Term enthalten:

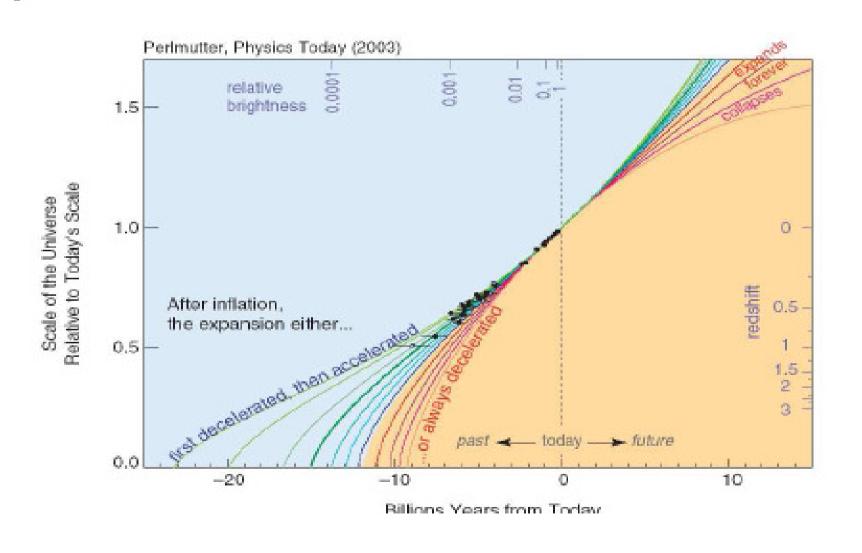
$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} + \Lambda g_{\mu\nu} = \frac{8\pi G_N}{c^4}T_{\mu\nu}$$

wobei die kosmologische Konstante  $\Lambda$  zeitlich konstant sein muss, damit die BIANCI-Identitäten erfüllt sind. Damit sehen dann die Gleichungen für  $\dot{R}$  und  $\ddot{R}$  folgendermaßen aus:

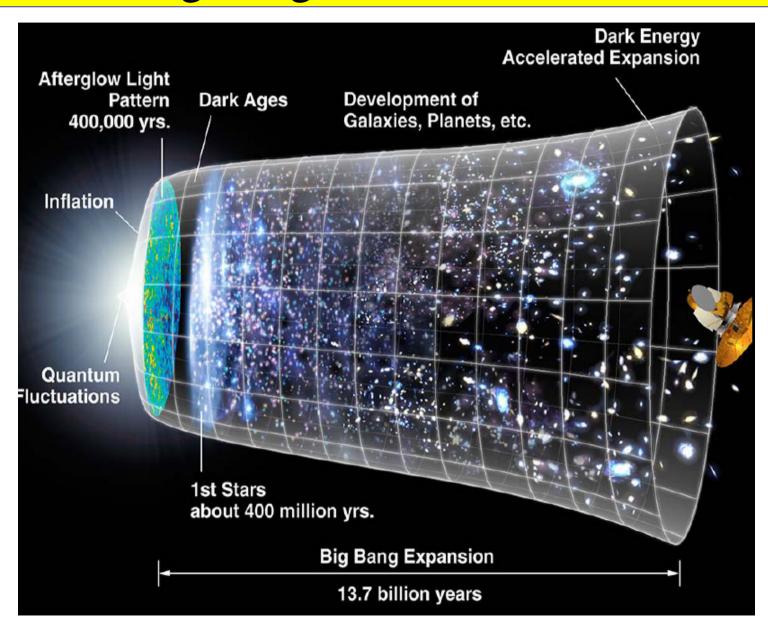
$$\left(\frac{\dot{R}(t)}{R(t)}\right)^{2} = \frac{8\pi G_{N}}{3}\rho(t) - \frac{Kc^{2}}{R^{2}(t)} + \frac{\Lambda c^{2}}{3} \tag{38}$$

$$\frac{\ddot{R}(t)}{R(t)} = -\frac{4\pi G_N}{3c^2} \left(3p(t) + \rho(t)c^2\right) + \frac{\Lambda c^2}{3}$$
(39)

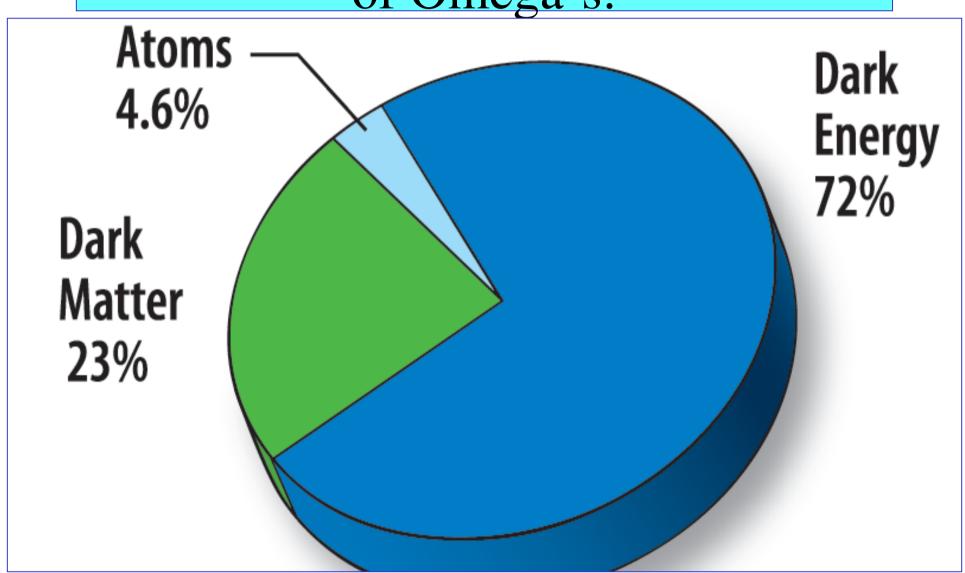
## Alternative forms of cosmic expansion:



#### The Big-bang Universe with K=0!:



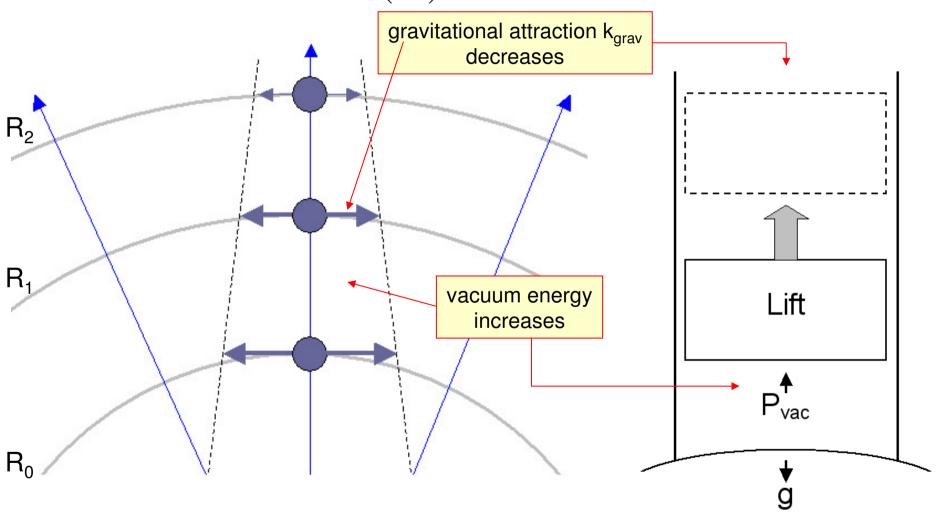
## What constitutes the world in terms of Omega's:



#### The problem with a constant vacuum energy density:

$$(\Lambda = const.)$$

$$k_{grav} = -\frac{8\pi G \rho_0 R_0}{3} \left(\frac{R_0}{R}\right)^2 \quad \text{= intermaterial gravitational force between co-moving masses}$$



#### The anthropic Lambda-miracle:

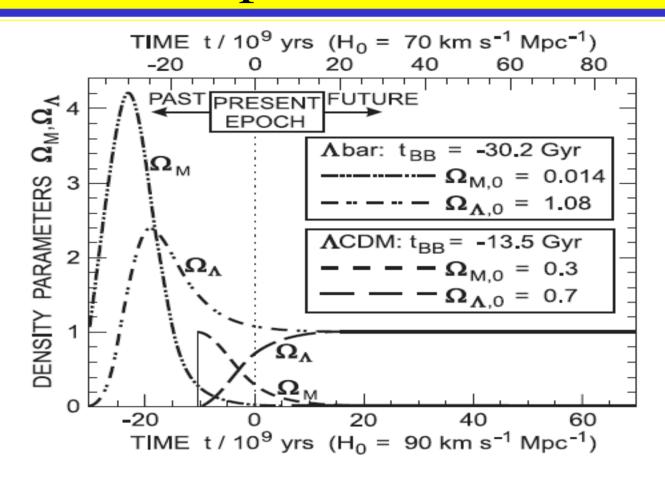
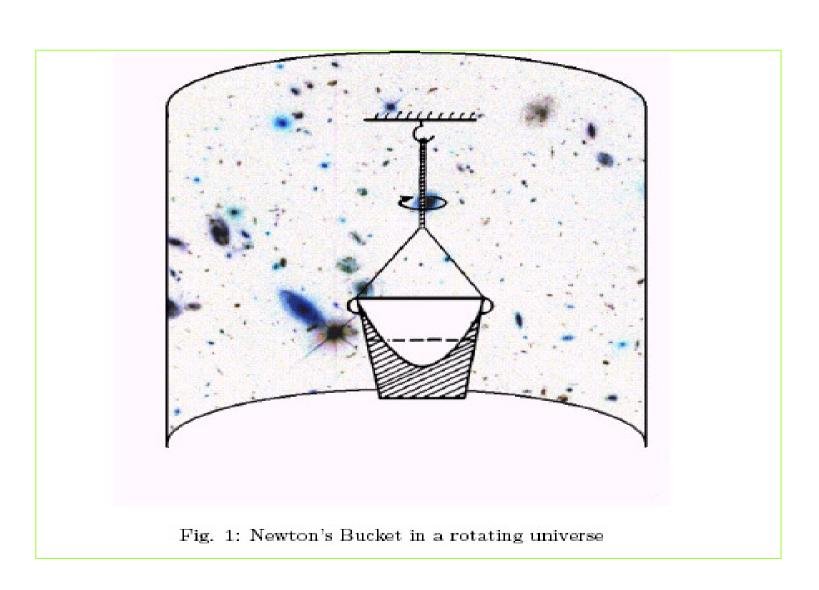


Fig. 7. Evolution of  $\Omega_{\rm M}$  and  $\Omega_{\Lambda}$  in the  $\Lambda$ bar and  $\Lambda$ CDM models (Models 1 and 6 in Figs. 2 and 3). Time is set to zero at the present epoch;  $t_{\rm BB}$ , the time of the big bang, is calculated using  $h_0 = 0.9$  for  $\Lambda$ bar (bottom scale) and  $h_0 = 0.7$  for  $\Lambda$ CDM (top scale). Compare Fig. 3.

### Do we have the physical concepts right in FLRW-cosmologies?

- Is the mass of the universe conserved?
- What is the effective cosmic density?
- How is gravitational binding energy entering the ART field equations?
- Is isotropic cosmic curvature reasonable?
- How all of this is reflected in changes of the vacuum energy density?

## What is the absolute reference system for centrifugal forces?



### The total universe as the cosmic reference system!

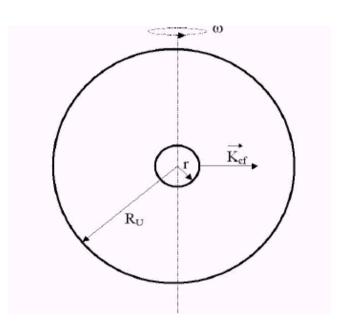


Fig 2.: Illustration of the earth at rest in a rotating universe

$$\overrightarrow{K}_{cf,a} = \psi \cdot \overrightarrow{K}_{cf,b}$$

 $\psi$  can be calculated as:

$$\psi = \frac{2GM_U}{c^2 R_U} = \frac{R_{S,U}}{R_U}$$

#### The instantaneous mass of the universe?

$$M_{\rm u}(t)c^2 = 4\pi\rho_0(t)c^2 \int_0^{R_{\rm u}} \frac{\exp(\lambda(r)/2)r^2 dr}{\sqrt{1 - (\frac{Hr}{c})^2}}$$

$$\exp[-\lambda(r)] = 1 - \frac{8\pi G}{rc^2} \rho_0 \int_0^r \gamma(x) x^2 dx$$

$$R_{
m u}=rac{1}{\pi}\sqrt{rac{c^2}{2G
ho_0}}$$

$$ho_0(R_{
m u}) = rac{c^2}{2\pi^2 G R_{
m u}^2}$$

#### Hoyle's creation theory

$$C_{\mu\nu} = \frac{\partial C_{\mu}}{\partial x^{\nu}} - \Gamma^{\alpha}_{\mu\nu} C_{\alpha} \qquad C_{\mu\nu} = -3R\dot{R}\frac{\partial_{\mu\nu}}{\partial A}$$

$$C_{\mu\nu} = -3R\dot{R}\frac{\delta_{\mu\nu}}{cA}$$

$$G_{\mu\nu} - \frac{1}{2}g_{\mu\nu}G + C_{\mu\nu} = \frac{8\pi\gamma}{c^4}T_{\mu\nu}$$

$$R = \exp[ct/A]$$

$$\dot{\rho}_{\rm H} = \frac{c}{A} \rho_{\rm H} = \frac{3c^5}{8\pi\gamma A^3} = \frac{c^5 A_{\rm H}^{3/2}}{8\pi\gamma\sqrt{3}}$$

### Vacuum energy and mass creation as analogous actions?

$$\Lambda^{3/2} = \frac{8\pi G\sqrt{3}}{c^5}\dot{\rho}$$

$$M_u = M_{u0} \exp\left[\frac{c(t-t_0)}{A}\right] = (\frac{M_{u0}}{R_0})R(t)$$

### A curved universe with metrical reactions to potential energy density (E.Fischer,1993):

$$T^p_{\mu\nu} = -C\frac{\rho_c}{\Gamma}g_{\mu\nu}$$

$$\frac{c^2k}{S^2} + \frac{\dot{S}^2}{S^2} + 2\frac{\ddot{S}}{S} = \frac{\kappa C\rho_{\rm c}}{S}$$

$$-3(\frac{c^{2}k}{S^{2}} + \frac{\dot{S}^{2}}{S^{2}}) = -\kappa\rho_{c} - \frac{\kappa C\rho_{c}}{S}$$

$$\frac{\ddot{S}}{S} = \frac{\kappa \rho_{\rm c}}{6} \left(\frac{S_0}{S} - 1\right)$$

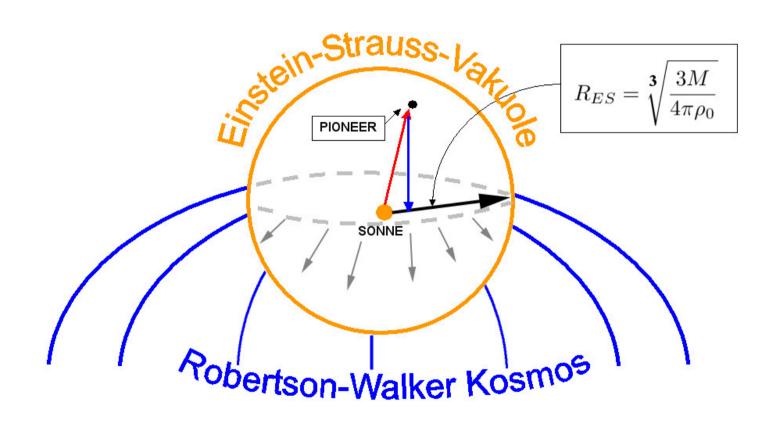
#### Curvature energy and mass creation:

$$\hat{T}_{00} = T_{00} + T_{00}^p = (\rho - C\frac{\rho}{\Gamma})g_{00}$$

$$\dot{\rho}^* = \frac{d}{dt} [\rho (1 - C \frac{1}{\Gamma})]$$

$$\dot{
ho}^* = 
ho C rac{1}{\Gamma^2} \dot{\Gamma}$$

#### Local masses in the global Universe?



### What is the effective cosmic mass density?

$$\rho^* = \frac{M}{V_{\rm ES}^3} = \frac{\frac{4\pi}{3}\rho_{\rm o}R_{\rm ES}^3}{V_{\rm ES}^3}$$

$$\rho^* = \frac{\frac{4\pi}{3}\rho_{\rm o}R_{\rm ES}^3}{4\pi(\frac{3c^2}{8\pi G\rho_{\rm o}})^{3/2}[\frac{1}{2}\arcsin\xi_{\rm ES} - \frac{\xi_{\rm ES}}{2}\sqrt{1-\xi_{\rm ES}^2}]}$$

$$\rho^* = \rho_0 \frac{1}{1 + \frac{3}{10} \xi_{ES}^2} \simeq \rho_0 \cdot (1 - \frac{3}{10} \Psi^2 \rho_0^{1/3})$$

## The effective cosmic mass density (high density limit):

$$ho^* = rac{M}{V_{
m ES}^3} = rac{rac{4\pi}{3}
ho_{
m o}R_{
m ES}^3}{V_{
m ES}^3}$$

$$\xi_{\rm ES} = \sqrt{\frac{8\pi G \rho_0}{3c^2}} R_{\rm ES} = \frac{\dot{S}}{cS} R_{\rm ES} = \frac{S_{\rm u} H}{c} \frac{R_{\rm ES}}{S_{\rm u}} = \frac{R_{\rm ES}}{S_{\rm u}} \ll 1$$

$$\rho^* = \rho_0 \frac{\xi_{\rm ES}^3}{\frac{3}{2} [\arcsin \xi_{\rm ES} - \xi_{\rm ES} \sqrt{1 - \xi_{\rm ES}^2}]}$$

### The local spacetime and the expansion of the

#### Einstein-Straus vacuole

$$r_{ES} = \left(\frac{3M}{4\pi\rho_0}\right)^{1/3}$$

$$\frac{\dot{r}_{ES}}{r_{ES}} = \frac{\dot{R}_0}{R_0} = H_0$$

#### The local world embedded in a vacuumenergy-loaded universe

$$c^2 \dot{M}(t) = -(4\pi R_{\rm ES}^2 \dot{R}_{\rm ES}) p_{\rm vac}$$

$$p_{\rm vac} = -\frac{3-n}{3}\rho_{\rm vac}c^2$$

$$c^2 \dot{M}(t) = 3M \frac{\dot{R}_{ES}}{R_{ES}} \frac{3 - n}{3} \frac{\rho_{\text{vac}}}{\rho_{\text{mat}}} c^2$$

$$\frac{\dot{M}(t)}{M} = 3 \frac{\dot{R}_{\rm ES}}{R_{\rm ES}} \frac{1}{3} \frac{\rho_{\rm vac}}{\rho_{\rm mat}}$$

#### Towards more realistic universes: The 2-phase structured universe

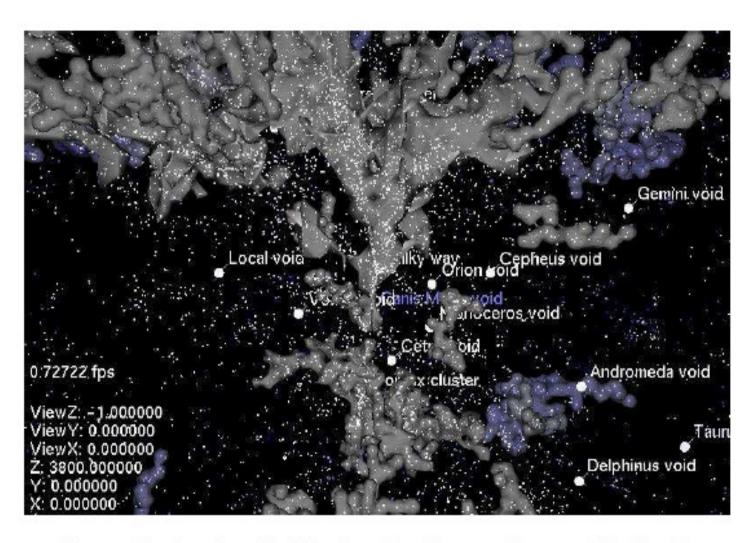
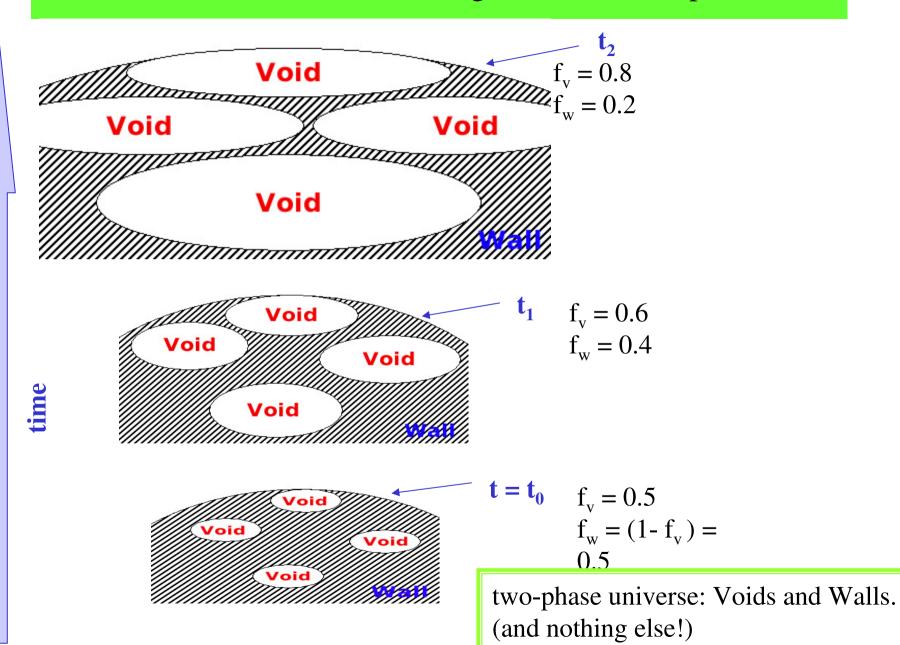


Fig. 1. Local voids and bubbles from the 6df survey. Courtesy of A. Fairall.

#### Voids and Walls: !Non-homologous structure expansions!



### General-Relativistic spacetime averages and GRT FRW equations

$$\langle \mathcal{R} \rangle \equiv \left( \int_{\mathcal{D}} \mathrm{d}^3 x \sqrt{\det {}^3\!g} \mathcal{R}(t, \mathbf{x}) \right) / \mathcal{V}(t)$$

with  $V(t) \equiv \int_{\mathcal{D}} d^3x \sqrt{\det {}^3g}$ . The important lesson of Buchert averaging is that time evolution and averaging to do not commute.<sup>5</sup> Generally for any scalar  $\Psi$ ,

$$\frac{\mathrm{d}}{\mathrm{d}t}\langle\Psi\rangle - \langle\frac{\mathrm{d}\Psi}{\mathrm{d}t}\rangle = \langle\Psi\vartheta\rangle - \langle\vartheta\rangle\langle\Psi\rangle \tag{1}$$

The fact that the r.h.s. of (1) does not vanish, as is the case for the FLRW cosmologies, is a manifestation of *backreaction*.

Applied to the equations of cosmic evolution one obtains the exact Buchert equations

$$3\frac{\dot{\bar{a}}^2}{\bar{a}^2} = 8\pi G\langle\rho\rangle - \frac{1}{2}\langle\mathcal{R}\rangle - \frac{1}{2}\mathcal{Q},\tag{2}$$

$$3\frac{\ddot{\bar{a}}}{\bar{a}} = -4\pi G \langle \rho \rangle + \mathcal{Q},\tag{3}$$

$$\partial_t \langle \rho \rangle + 3 \frac{\dot{a}}{\bar{a}} \langle \rho \rangle = 0,$$
 (4)

#### The back-reaction of curvature averages

$$\mathcal{Q} \equiv \frac{2}{3} \left( \langle \vartheta^2 \rangle - \langle \vartheta \rangle^2 \right) - 2 \langle \sigma \rangle^2$$

$$\bar{\Omega}_{M} = \frac{8\pi G \bar{\rho}_{M0} \bar{a}_{0}^{3}}{3\bar{H}^{2} \bar{a}^{3}} \; ; \; \bar{\Omega}_{k} = \frac{-k_{\rm v} f_{\rm vi}^{2/3} f_{\rm v}^{1/3}}{\bar{a}^{2} \bar{H}^{2}} \; ; \; \bar{\Omega}_{\mathcal{Q}} = \frac{-\dot{f_{\rm v}}^{2}}{9 f_{\rm v} (1 - f_{\rm v}) \bar{H}^{2}}$$

$$f_{\rm v} = \frac{3f_{\rm v0}\bar{H}_{\rm 0}t}{3f_{\rm v0}\bar{H}_{\rm 0}t + (1 - f_{\rm v0})(2 + f_{\rm v0})}$$

$$q = \frac{-(1 - f_{\rm v}) (8f_{\rm v}^3 + 39f_{\rm v}^2 - 12f_{\rm v} - 8)}{(4 + f_{\rm v} + 4f_{\rm v}^2)^2}$$

## Redshift-magnitude relation in different universes

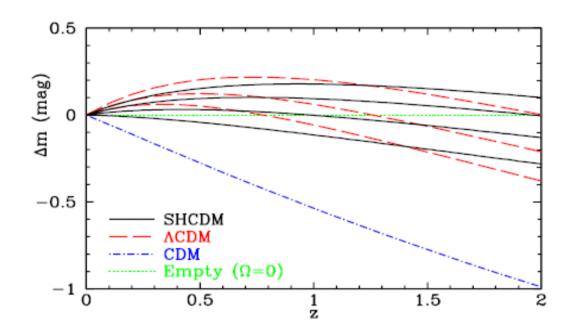


FIG. 1: The apparent magnitude difference as a function of redshift between SHCDM and  $\Lambda$ CDM models compared to a model empty universe. The SHCDM models from top to bottom are for  $\Psi_{\ell 0} = -1.0$ , -0.75, -0.5, and -0.25, while the  $\Lambda$ CDM models from top to bottom are for  $\Omega_{\Lambda} = 0.8$ , 0.7, and 0.6 (all with w = -1). Also indicated is the CDM model ( $\Omega_{\Lambda} = 0$ ).

### Structured cosmic matter in a 2-phase universe as analogy to vacuum energy

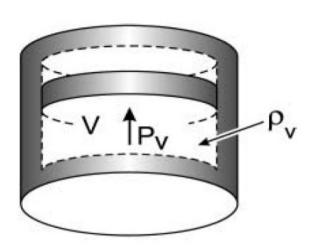
$$\bar{\rho}_2 = \rho_v f_v + \rho_w f_w = \rho_v f_v + \rho_w (1 - f_v)$$

$$\rho_{vac} = \frac{\bar{\rho}_2(1 - 2\bar{q}_2)}{2(\bar{q}_2 + 1)}$$

$$\bar{\rho}_2(f_v \ge f_{vc}) = \bar{\rho}_2 - \bar{\rho}_{vac}(f_v \ge f_{vc}) = \bar{\rho}_2(1 - \frac{1 - 2\bar{q}_2}{2(\bar{q}_2 + 1)})$$

$$f_v \ge f_{vc} = 0.57$$

#### Vacuum action in GRT?



**Fig. 6** Equation of state of the vacuum. As the vacuum does work to push out the piston, it creates *more* vacuum inside the chamber, increasing its internal energy. Ordinary intuition fails here if we imagine the piston in a typical laboratory environment, surrounded by high-pressure gas. In cosmology, there is no container, and no "outside" at all

$$dU = -p_{\text{Vac}}dV$$

$$dU=\rho_{\text{Vac}} c^2 dV$$

$$p_{\text{Vac}} = -\rho_{\text{Vac}}c^2$$

$$T_{\mu\nu}^{\text{Vac}} = (\rho_{\text{Vac}}c^2 + p_{\text{Vac}})U_{\mu}U_{\nu} - p_{\text{Vac}}g_{\mu\nu}$$
  
=  $\rho_{\text{Vac}}c^2g_{\mu\nu}$ 

#### A rational concept of empty space

$$S_{\rm GR} = \frac{-1}{16\pi G} \int d^4x \sqrt{-g} \left(R + 2\Lambda\right) + S_{\rm Mat}$$

$$R_{\mu\nu} - \frac{1}{2}R g_{\mu\nu} = \frac{-8\pi G}{c^4} \left( T_{\mu\nu}^{\text{Mat}} + \frac{\Lambda c^4}{8\pi G} g_{\mu\nu} \right)$$

#### the spacetime of "nothing

$$R_{\mu
u} - rac{1}{2} R \, g_{\mu
u} = - \Lambda_{
m Eff} g_{\mu
u},$$

where the *effective* cosmological constant is

$$\Lambda_{
m Eff} \equiv \Lambda + rac{8\pi G 
ho_{
m Vac}}{c^2},$$

#### !No photon redshift in empty space!

$$ds^{2} = c^{2}dt^{2} - R^{2} \left[ \frac{dr^{2}}{(1 - kr^{2})} + r^{2} \left( d\theta^{2} + \sin^{2}\theta d\phi^{2} \right) \right]$$

$$R(t) \approx \begin{cases} \cosh\left(\sqrt{\Lambda_{\rm Eff}/3} \, ct\right) & \text{if} \quad k = +1 \\ \exp\left(\sqrt{\Lambda_{\rm Eff}/3} \, ct\right) & \text{if} \quad k = 0 \\ \sinh\left(\sqrt{\Lambda_{\rm Eff}/3} \, ct\right) & \text{if} \quad k = -1 \end{cases}$$

A photon will thus experience conting interaction with *empty space* at a rate

$$\frac{dz}{dt} = \sqrt{\frac{\Lambda_{\rm Eff}}{3}}c,$$

where  $t \equiv t_{\rm O} - t_{\rm E}$ , unless we set  $\Lambda_{\rm Eff} = 0$ 

$$\Lambda_{eff} = \frac{8\pi G}{c^2} (\rho_{vac} - \rho_{vac,0})$$

#### The "zero energy" universe:

$$E = \int_{0}^{V^{3}} (\rho c^{2} + 3p) \sqrt{-g_{3}} d^{3}V = \frac{4\pi}{3} S^{3} (\rho c^{2} + 3p)$$

$$\rho = \rho_b + \rho_d + \rho_{vac}$$

and the total pressure is given by

$$p = p_b + p_d + p_{\text{vac}}$$

In the present phase of the evolution of the universe, baryonic and dark matter can be considered as cold and pressure-less, i.e.,  $p_b + p_d = 0$ . Assuming, furthermore, a general dependence of  $\rho_{\text{vac}} \sim S^{-n}$  (see Fahr and Heyl 2006b), one then obtains  $p = p_{\text{vac}} = -\frac{3-n}{3}\rho_{\text{vac}}c^2$  and finds

$$E = \frac{4\pi}{3}S^3c^2(\rho_b + \rho_d + (n-2)\rho_{\text{vac}})$$
 (15)

$$\frac{3c^2}{2\pi GS^2} = (\rho_b + \rho_d + (n-2)\rho_{\text{vac}})$$

Now, the requirement that the total energy of the universe L = E + U vanishes – with E and U given by

$$U = -\frac{8\pi^2 G}{15} (\rho_b + \rho_d + (n-2)\rho_{\text{vac}})^2 S^5$$

$$\Phi(r) = -\frac{2}{3}\pi G(\rho_{b} + \rho_{d} + (n-2)\rho_{vac})r^{2}$$

$$U = \int_0^S 4\pi r^2 (\rho_b + \rho_d + (n-2)\rho_{\text{vac}}) \Phi(r) dr$$

#### The "Zero-Energy" universe

### We could get the physical concepts right,...if....!

- Is the mass of the universe conserved?
- There is mass creation due to vacuum decay!
- What is the effective cosmic density?
- It is the metrically modulated proper density!
- How is gravitational binding energy entering the ART field equations?
- It reduces the effective density!
- How changes the vacuum energy density?
- It decays inversely proportional to the square of S

# All stories and books have a begin; the universe has none! unless we make a story out of it!



## But all over the world doubts come up.....

